THIN MATERIALS WITH INDUCED METRIC

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- An old story: dimension reduction 3d → 2d, many papers in linearized elasticity, nonlinear elasticity with challenging questions,
- More recent: same idea with a different metric.

"A material aims at achieving a prescribed metric." Is this possible at the 3d-level? What are the induced 2d-models?

Part 1. Classical results, 1993-2006

1.1. 3d-elastic body classical internal elastic energy

$$\begin{split} I(\Phi) = \int_{\Omega} W(\nabla \Phi(x)) \, dx, \quad W: \ \mathbb{M}^{3+} \mapsto \mathbb{R}^+ \ \text{stored energy density,} \\ \mathbb{M}^{3+} := \{F \in \mathbb{M}^3; \det F > 0\}. \end{split}$$

Hypotheses

- Frame indifference $W(F) = \tilde{W}(F^T F)$,
- ► $W(Id) = 0; W(F) = 0 \Leftrightarrow F \in SO(3).$ Then, W min on $SO(3), \frac{\partial W}{\partial F}(Id) = 0, \Omega$ natural state.

Under external loads, an equilibrium state in a "minimizer" on $W^{1,p}(\Omega;\mathbb{R}^3)$ of

$$I(\Phi) - \int_{\Omega} f \cdot \Phi \, dx$$
 with b.c.

What are the minimizers when f = 0? rigid motions (b.c. compatible) Indeed,

$$I(\Phi) = \int_{\Omega} W(\nabla \Phi(x)) \, dx, \quad I(\Phi) = 0 \text{ iff } \nabla \Phi(x)^T \nabla \Phi(x) = Id,$$

and Liouville theorem

 $[\Phi:\Omega\mapsto\mathbb{R}^3,\nabla\Phi(x)^T\nabla\Phi(x)=Id,\det\nabla\Phi(x)>0]\Leftrightarrow[\exists R\in SO(3),\forall x,\Phi(x)=Rx+c].$

Three proofs

- 1. $\psi = \Phi \Phi(0)$, ψ preserves the inner product, hence is linear,
- 2. differentiate $\partial_i \Phi(x) \cdot \partial_j \Phi(x) = \delta_{ij}$, obtain $\partial_{ij} \Phi(x) = 0$,
- scheme for more sophisticated results: ∇Φ(x) ∈ SO(3), hence ∇Φ = Cof ∇Φ; ΔΦ = div ∇Φ = div Cof ∇Φ = 0 (Piola). Thus, Φ is harmonic and C[∞]. Now, |∇Φ(x)|² = tr(∇Φ(x))^T∇Φ(x) = 3,

$$0 = \Delta(|\nabla \Phi(x)|^2) = 2|\nabla^2 \Phi(x)|^2 + \nabla \Phi(x) : \Delta \nabla \Phi(x) = 2|\nabla^2 \Phi(x)|^2.$$

Order 2 derivatives are null.

1.2. Thin model hierarchy

Find the limit behavior of (almost) minimizers $\Phi^h: \Omega^h = \omega \times]-h, h[\mapsto \mathbb{R}^3$ of

$$\frac{1}{2h}\int_{\Omega^h}W(\nabla\Phi(x))\,dx-\int_{\Omega^h}f^h\cdot\Phi\,dx\quad\text{with b.c.}$$

- hierarchy of four models by asymptotic exp., with no a priori assumption, Fox, R. & Simo, ARMA, 1993
 - 1. nonlinear membrane model
 - 2. (inextensional) nonlinear bending model
 - 3. von Kármán model
 - 4. linear elasticity model

driven by the loading magnitude, can be expressed in terms of the internal energy magnitude.

- ▶ rigorous convergence for the membrane model, Le Dret & R., JMPA, 1995
- rigorous convergence for the bending model, Friesecke, James & Müller, CPAM, 2002
- rigorous convergence for the vK model, Friesecke, James & Müller, ARMA, 2006; R., 2002.

Part 2. Induced metric

2.1. 3d-body with new internal energy, growth-induced

Let be given a

metric $G: x \in \Omega \mapsto G(x) \in \mathbb{S}^{3+} = \{\text{positive definite symetric}\}.$

The material aims at satisfying

$$abla \Phi^T
abla \Phi = G, \, \det(
abla \Phi(x)) > 0.$$

Modeling proposed by Kupferman, Sharon, circa 2008.

Alternative formulation: let $A^2(x) = G(x)$, $A \in \mathbb{S}^{3+}$. There is an internal energy that reads

$$I(\Phi) = \int_{\Omega} W(\nabla \Phi(x) A^{-1}(x)) dx, \quad \det(\nabla \Phi(x)) > 0$$

still with W = 0 on SO(3). When is this energy 0?

$$I(\Phi) = 0 \quad \text{when} \quad \nabla \Phi(x) A^{-1}(x) \in SO(3), \det(\nabla \Phi(x)) > 0,$$

i.e.,
$$\nabla \Phi^T \nabla \Phi = G, \det(\nabla \Phi(x)) > 0.$$

Is $\nabla \Phi^T \nabla \Phi = G$, det $(\nabla \Phi(x)) > 0$ possible?

"Easy" : use $\partial_{ik}\partial_j \Phi = \partial_{ij}\partial_k \Phi$ in $\nabla \Phi^T \nabla \Phi = G$. Obtain $\mathscr{R} = 0$ where $\mathscr{R}_{aiik} = \partial_i \Gamma_{ikq} - \partial_k \Gamma_{ijq} + \Gamma^{\rho}_{ij} \Gamma_{kqp} - \Gamma^{\rho}_{ik} \Gamma_{jqp}$, six "independent" nonzero entries,

$$2\Gamma_{ijq} = \partial_j g_{iq} + \partial_i g_{jq} - \partial_q g_{ij}, \Gamma^p_{ij} = g^{pq} \Gamma_{ijq}, (g^{pq}) = G^{-1}.$$

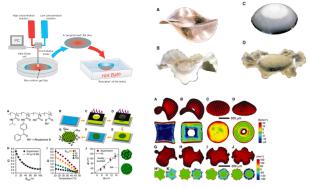
Conversely,

fundamental theorem of Riemannian geometry: if $\mathscr{R} = 0$, there exists Φ , det $\nabla \Phi(x) > 0$ and $\nabla \Phi^T \nabla \Phi = G$, G said flat metric.

Additional result, Lewicka & Pakzad, 2011 if $\mathscr{R} \neq 0$, we know that $\nexists \Phi \in H^1$, $||\operatorname{dist}(\nabla \Phi(\cdot), SO(3)A(\cdot))||_{L^2(\Omega)} = 0$. In fact,

$\inf_{\Phi \in H^{\mathbf{1}}} ||\operatorname{dist} (\nabla \Phi(\cdot), SO(3)A(\cdot))||_{L^{2}(\Omega)} > 0.$

Not obvious, because there is no reason why the inf should be attained. Method of proof: reminiscent of the quantitative rigidity estimate. 2.2. Induced thin models, of interest even in the absence of loads or b.c.



Problem setting: behavior of (almost) minimizers Φ^h of

$$I^{h}(\Phi) = \frac{1}{2h} \int_{\Omega^{h}} W(\nabla \Phi(x) A^{-1}(\bar{x})) \, dx, \ \Phi : \Omega^{h} = \omega \times] - h, h[\mapsto \mathbb{R}^{3},$$

A only depends on $\bar{x} = (x_1, x_2)$. Magnitude of $\inf I^h$? $I^h(\Phi)$ reads as well

$$I^{h}(\Phi) = \int_{\Omega} W(\nabla_{h} \Phi(x) A^{-1}(\bar{x})) dx, \quad \Phi : \Omega = \omega \times] - 1, 1[\mapsto \mathbb{R}^{3},$$

 $\nabla_h \Phi = (\partial_1 \Phi, \partial_2 \Phi, \frac{1}{h} \partial_3 \Phi).$

where we keep the same notation $\boldsymbol{\Phi}$

$$I^{h}(\Phi) = \int_{\Omega} W(\nabla_{h} \Phi(x) A^{-1}(\bar{x})) \, dx, \quad \nabla_{h} \Phi = (\partial_{1} \Phi, \partial_{2} \Phi, \frac{1}{h} \partial_{3} \Phi)$$

Order 0 model: Generalized membrane energy

Usual case A = Id. For $\overline{F} \in \mathbb{M}_{3 \times 2}$, let $W_0(\overline{F}) = \min\{W(F); F \in \mathbb{M}^3, F_{3 \times 2} = \overline{F}\}$.

$$\mathcal{I}^{h} \xrightarrow{\Gamma - L^{p}(\Omega)} \mathcal{I}_{0}, \ \mathcal{I}_{0}(\Phi) = \begin{cases} \int_{\omega} QW_{0}(\bar{\nabla}\varphi(\bar{x})) d\bar{x}, \ \Phi = \varphi \in W^{1,p}(\omega; \mathbb{R}^{3}), \\ +\infty, \ \Phi \in L^{p}(\Omega; \mathbb{R}^{3}) \setminus W^{1,p}(\omega; \mathbb{R}^{3}). \end{cases}$$

Easily,

 $QW_0(\bar{\nabla}\varphi(\bar{x}))$ only depends on $(\nabla\varphi(\bar{x}))^T \nabla\varphi(\bar{x})$ metric tensor $l_0(\varphi) = 0$ for $\varphi: \omega \mapsto \mathbb{R}^3$ Eucl. isometry $\partial_\alpha \varphi \cdot \partial_\beta \varphi = \delta_{\alpha\beta}$ (and more).

Note that zero-energy deformations for the 3d-model are SO(3), 2d-limit has much more.

Fox, R. & Simo, Le Dret & R., Ben Belgacem.

General $A(\bar{x})$. $\bar{F} \in \mathbb{M}_{3 \times 2}$, $W_0(\bar{x}, \bar{F}) := \min\{W(FA^{-1}(\bar{x})); F \in \mathbb{M}^3, F_{3 \times 2} = \bar{F}\}.$

$$I^{h} \xrightarrow{\Gamma - L^{p}(\Omega)} I_{0}, I_{0}(\Phi) = \begin{cases} \int_{\omega} QW_{0}(\bar{x}, \bar{\nabla}\varphi(\bar{x})) d\bar{x}, \Phi = \varphi \in W^{1,p}(\omega; \mathbb{R}^{3}), \\ +\infty, \Phi \in L^{p}(\Omega; \mathbb{R}^{3}) \setminus W^{1,p}(\omega; \mathbb{R}^{3}). \end{cases}$$

Now, $l_0(\varphi) = 0$ as soon as

$$\exists b(\bar{x}) \in \mathbb{R}^3 \, \text{s.t} \left(\partial_1 \varphi(\bar{x}) | \partial_2 \varphi(\bar{x}) | b(\bar{x}) \right) A^{-1}(\bar{x}) \in SO(3)$$

i.e, omitting \bar{x} ,

$$\begin{pmatrix} \bar{\nabla} \phi^{\mathsf{T}} \bar{\nabla} \phi & \partial_{\alpha} \phi \cdot b \\ \partial_{\alpha} \phi \cdot b & |b|^2 \end{pmatrix} = \begin{pmatrix} \mathsf{G}_{\alpha\beta} & \mathsf{G}_{\alpha3} \\ \mathsf{G}_{\alpha3} & \mathsf{G}_{33} \end{pmatrix}, \, \mathsf{det} \left(\partial_1 \phi | \partial_2 \phi | b \right) > 0.$$

It suffices to satisfy

$$\bar{\nabla}\varphi^T\bar{\nabla}\varphi = [G_{\alpha\beta}]_{\alpha,\beta=1,2} := G_{2\times 2}.$$
 (1)

Indeed, b follows: given components along two vectors, norm and orientation.

Is (1) realizable? Yes, Nash-Kuiper *circa* 1954, with C^1 -regularity, not C^2 (curvature obstruction, Hilbert counter-example, 1901).

Even if the 3d-model has no zero-energy deformations (and inf > 0), 2d-limit has.

Footnote: Isometric immersion of the flat torus into \mathbb{R}^3 , based on Gromov construction (plenty of cheap solutions), Hevea project.



Order 2 model: Generalized bending energy

Usual case A = Id. For $F^{\sharp} \in \mathbb{M}_{2 \times 2}$, let $W_2(F^{\sharp}) = \min\{W''(Id)(F,F); F \in \mathbb{M}^3, F_{2 \times 2} = F^{\sharp}\},\$ $= 2\mu |\frac{F^{\sharp} + F^{\sharp}T}{2}|^2 + \frac{2\mu}{2\mu + \lambda} (\operatorname{tr} F^{\sharp})^2.$

$$\frac{I^{h}}{h^{2}} \xrightarrow{\Gamma - H^{1}(\Omega)} I_{2}, I_{2}(\Phi) = \begin{cases} \frac{1}{3!} \int_{\omega} W_{2}\left((\bar{\nabla} \varphi^{T} \bar{\nabla} n)(\bar{x})\right) d\bar{x}, \ \Phi = \varphi \in H^{2}(\omega; \mathbb{R}^{3}), iso, \\ +\infty \text{ otherwise.} \end{cases}$$

iso: $|\partial_1 \varphi| = 1, |\partial_2 \varphi| = 1, \partial_1 \varphi \cdot \partial_2 \varphi = 0, \quad \overline{\nabla} \phi^T \overline{\nabla} n$: surface curvature tensor Obviously,

 $l_2(\varphi) = 0$ for $\varphi : \omega \mapsto \mathbb{R}^3$ Eucl. isometry and null curvature tensor (first and second forms equal to 0): $\varphi \in O(3)(\bar{x}, 0)$.

Fox, R. & Simo, Friesecke, James & Müller, Pantz

Extended
$$A(\bar{x}) = \begin{pmatrix} A_{\alpha\beta}(\bar{x}) & 0\\ 0 & 0 & 1 \end{pmatrix}$$
.
 $F^{\sharp} \in \mathbb{M}_{2 \times 2}, \ W_2(\bar{x}, F^{\sharp}) = \min\{W''(Id)(A^{-1}FA^{-1})^{(2)}; F \in \mathbb{M}^3, F_{2 \times 2} = F^{\sharp}\}.$
 $\frac{I^h}{h^2} \xrightarrow{\Gamma - H^1(\Omega)} I_2, \ I_2(\Phi) = \begin{cases} \frac{1}{3!} \int_{\omega} W_2(\bar{x}, ((\bar{\nabla}\varphi)^T \bar{\nabla}n)(\bar{x})) d\bar{x}, \ \Phi = \varphi \in H^2(\omega; \mathbb{R}^3), iso, \\ +\infty \text{ otherwise.} \end{cases}$

iso: $(\overline{\nabla} \varphi)^T \overline{\nabla} \varphi = G_{\mathbf{2} \times \mathbf{2}}$ Lewicka& Pakzad

General
$$A(\bar{x})$$
.
 $W_2(\bar{x}, F^{\sharp}) = \min\{W''(Id)(A^{-1}(\bar{x})FA^{-1}(\bar{x}))^{(2)}; F \in \mathbb{M}^3, F_{2\times 2} = F^{\sharp}\}.$
 $\frac{I^h}{h^2} \xrightarrow{\Gamma - H^1(\Omega)} I_2, I_2(\Phi) = \begin{cases} \frac{1}{3!} \int_{\omega} W_2(\bar{x}, ((\bar{\nabla} \varphi^T \bar{\nabla} \mathbf{b})(\bar{x})) d\bar{x}, \Phi = \varphi \in H^2(\omega; \mathbb{R}^3), iso Herwise. \end{cases}$

iso: $(\bar{\nabla}\varphi)^T \bar{\nabla}\varphi = G_{2\times 2}$ where $b \in H^1 \cap L^{\infty}$ uniquely defined in terms of φ as in slides 8 and 9 *i.e.*

$$\begin{pmatrix} \bar{\nabla} \varphi^T \bar{\nabla} \varphi & \partial_{\alpha} \varphi \cdot b \\ \partial_{\alpha} \varphi \cdot b & |b|^2 \end{pmatrix} = \begin{pmatrix} G_{\alpha\beta} & G_{\alpha3} \\ G_{\alpha3} & G_{33} \end{pmatrix}, \det(\partial_1 \varphi | \partial_2 \varphi | b) > 0.$$

Or, letting
$$Q = \left(\partial_1 \varphi(\bar{x}) | \partial_1 \varphi(\bar{x}) | b(\bar{x}) \right)$$
, there holds $Q^T Q = G$, det $Q > 0$.

Bhattacharya, Lewicka & Schäffner

Method of proof: extension of the quantitative rigidity estimate on slender domains, F.J.M, 2002 $\Omega \in \mathbb{R}^3$ given: $\exists C(\Omega) > 0$,

$$\forall \Phi \in H^{1}(\Omega; \mathbb{R}^{3}), \exists R \in SO(3), ||\nabla \Phi - R||_{L^{2}(\Omega)} \leq C(\Omega) ||\operatorname{dist}(\nabla \Phi, SO(3))||_{L^{2}(\Omega)}$$
 indep. *

 $\Omega^{h} = \omega \times] - h, h[\text{ or alternatively, } \nabla_{h} \text{ on } \Omega: \text{ roughly speaking, } \exists c(\omega) > 0,$

$$\forall h, \forall \Phi \in H^{1}(\Omega; \mathbb{R}^{3}), \exists R : \boldsymbol{\omega} \mapsto SO(3), \begin{cases} ||\nabla_{h} \Phi - R||_{L^{2}(\Omega)} \leq c||\operatorname{dist}(\nabla_{h} \Phi, SO(3))||_{L^{2}(\Omega)} \\ ||\nabla R||_{L^{2}(\boldsymbol{\omega})} \leq \frac{c}{h}||\operatorname{dist}(\nabla_{h} \Phi, SO(3))||_{L^{2}(\Omega)} \end{cases}$$

(bis)

$$\frac{I^{h}}{h^{2}} \xrightarrow{\Gamma - H^{1}(\Omega)} I_{2}, I_{2}(\Phi) = \begin{cases} \frac{1}{3!} \int_{\omega} W_{2}\left(\bar{x}, (\bar{\nabla}\varphi^{T}\bar{\nabla}b)(\bar{x})\right) d\bar{x}, \ \Phi = \varphi \in H^{2}(\omega; \mathbb{R}^{3}), iso, \\ +\infty \text{ otherwise.} \end{cases}$$

If min $I_2 = 0$, further information may be sought for.

$$\min I_2 = 0 \Leftrightarrow \exists \varphi \in H^2(\omega; \mathbb{R}^3), \, \bar{\nabla} \varphi(\bar{x})^T \bar{\nabla} \varphi(\bar{x}) = G_{2 \times 2}, \, \bar{\nabla} \varphi^T \bar{\nabla} b \text{ skew} \\ \Leftrightarrow \mathscr{R}_{1212} = \mathscr{R}_{1213} = \mathscr{R}_{1223} = 0 \, (\neq [\mathscr{R} = 0], \inf I^h > 0 \text{ allowed}).$$

Such φ is unique, because its 2nd fundamental form, in addition to its first fundamental form, is then given in terms of *G*.

Order 4 model: Generalized von Kármán enegy

Usual case A = Id. Recall that

 $\frac{l^{h}}{h^{2}} \xrightarrow{\Gamma - H^{1}(\Omega)} l_{2}, \ l_{2}(\varphi) = [\text{second fundamental form }]^{2} \text{ on iso}(\omega; \mathbb{R}^{3}),$

= 0 for $\varphi = R(\bar{x}, 0)$, $R \in O(3)$

Examine smaller orders of magnitude of I^h corresponding intuitively to

$${}^{``}\Phi^{h}(x) = (\bar{x}, 0) + h(u_{1}^{1}, u_{2}^{1}, x_{3} + u_{3}^{1}) + h^{2}\Phi^{2} + \cdots ."$$
Obtain $u_{1}^{1} = u_{2}^{1} = 0$ (through $\partial_{\alpha}u_{\beta}^{1} + \partial_{\beta}u_{\alpha}^{1} = 0$), $\Phi^{2} = u^{2} - x_{3}\partial_{\alpha}u_{3}^{1}e_{\alpha}$,
 $\frac{l^{h}}{h^{4}} \rightarrow l_{4}, l_{4}(u_{1}^{2}, u_{2}^{2}, u_{3}^{1}) = \int_{\omega} \left(2W_{2}\left(\frac{\partial_{\alpha}u_{\beta}^{2} + \partial_{\beta}u_{\alpha}^{2} + \partial_{\alpha}u_{3}^{1}\partial_{\beta}u_{3}^{1}}{2}\right) + \frac{1}{3!}W_{2}(\partial_{\alpha\beta}u_{3}^{1})\right) d\bar{x}$,

recall that W_2 is quadratic in its argument.

Fox, R. & Simo, Friesecke, James & Müller, R.

General $A(\bar{x})$. Lewicka, Raoult & Ricciotti. Start from $\min I_2 = 0$, *i.e.* $\Re_{1212} = \Re_{1213} = \Re_{1223} = 0$, *i.e.* $\exists \varphi \in H^2(\omega; \mathbb{R}^3)$, $\bar{\nabla} \varphi^T \bar{\nabla} \varphi = G_{2 \times 2}$ and $\bar{\nabla} \varphi^T \bar{\nabla} b$ skew, where $b: \omega \mapsto \mathbb{R}^3$ defined by $Q = (\partial_1 \varphi(\bar{x}) | \partial_1 \varphi(\bar{x}) | b(\bar{x})), Q^T Q = G, \det Q > 0$, or else $b = -(G^{33})^{-1}(G^{\alpha 3}\partial_{\alpha}\varphi) + (G^{33})^{-1/2}n$. First finding. Then $\inf I^h$ is indeed smaller: $\inf I^h \leq Ch^4$. Let simply $\Phi^h(\bar{x}, x_3) = \varphi(\bar{x}) + x_3 b(\bar{x}) + \frac{x_3^2}{2} d(\bar{x})$ (indep. of h). Which d? $\nabla \Phi^h(\bar{x}, x_3) = Q(\bar{x}) + x_3 B(\bar{x}) + \frac{x_3^2}{2} D(\bar{x})$,

with

$$Q = [\partial_{\alpha} \varphi, b], B = [\partial_{\alpha} b, d], D = [\partial_{\alpha} d, 0].$$

$$\begin{aligned} \nabla \Phi^{h} A^{-1}(\bar{x}, x_{3}) &= Q A^{-1}(\bar{x}) + x_{3} B A^{-1}(\bar{x}) + \frac{x_{3}^{2}}{2} D A^{-1}(\bar{x}) \\ &= (Q A^{-1}) (\mathsf{Id} + x_{3} A^{-1} Q^{T} B A^{-1} + x_{3}^{2} T). \\ W(\nabla \Phi^{h} A^{-1}) &= W (\mathsf{Id} + x_{3} A^{-1} Q^{T} B A^{-1} + x_{3}^{2} T). \end{aligned}$$

Make $Q^T B$ skew-symmetric (will kill the x_3^2 term in W''(Id)).

$$Q = [\partial_{\alpha} \varphi, b], B = [\partial_{\alpha} b, d], Q^{\mathsf{T}} B = \begin{pmatrix} \overline{\nabla} \varphi^{\mathsf{T}} \overline{\nabla} b & \overline{\nabla} \varphi^{\mathsf{T}} d \\ b^{\mathsf{T}} \overline{\nabla} b & b \cdot d \end{pmatrix},$$

b is already determined, we know that $\nabla \phi^T \nabla b$ is skew, then choose *d* s.t. $Q^T B$ skew: $Q^T d = (-b \cdot \partial_1 b, -b \cdot \partial_2 b, 0)^T$. Limit model. We already know that $\Phi^h \xrightarrow{h^1} \phi$, $\frac{1}{h} \partial_3 \Phi^h \xrightarrow{L^2} b$. Now,

$$U^{h}(\bar{x}) := \frac{1}{2h} \int_{-1}^{1} \left(\Phi^{h} - \left(\varphi + hx_{3}b \right) \right) dx_{3} \xrightarrow{H^{1}} u^{1}, \text{ sym} \left(\bar{\nabla} \varphi^{T} \bar{\nabla} u^{1} \right) = 0,$$

$$rac{1}{h}\operatorname{\mathsf{sym}}ig(ar
abla arphi^{ op} ar
abla^{ op} U^hig) o e^2 \in L^2(\omega;\mathbb{S}_2),$$

analog of
$$\partial_{\alpha} u_{\beta}^{1} + \partial_{\beta} u_{\alpha}^{1} = 0$$

analog of $\frac{\partial_{\alpha} u_{\beta}^{2} + \partial_{\beta} u_{\alpha}^{2}}{2}$

Limit energy given by

$$I_{4}(u^{1}, e^{2}) = I_{\nu \mathcal{K}}(u^{1}, e^{2}) = 2 \int_{\omega} W_{2}\left(\bar{x}, e^{2} + \frac{1}{2}(\bar{\nabla}u^{1})^{T}\bar{\nabla}u^{1} + \frac{1}{4!}\bar{\nabla}b^{T}\bar{\nabla}b\right)$$
$$+ \frac{1}{3!}\int_{\omega} W_{2}\left(\bar{x}, \bar{\nabla}\varphi^{T}\bar{\nabla}\rho^{1} + (\bar{\nabla}u^{1})^{T}\bar{\nabla}b\right)$$
$$+ \frac{2}{6!}\int_{\omega} W_{2}\left(\bar{x}, \operatorname{sym}(\bar{\nabla}\varphi^{T}\bar{\nabla}d) + \bar{\nabla}b^{T}\bar{\nabla}b\right)$$

where p^1 defined by $Q^T p^1 = (-b \cdot \partial_1 u^1, -b \cdot \partial_2 u^1, 0)^T_{\cdot}$ usual case $p^1 = (-\partial_{\alpha} u^1_3, 0)$ Remark: the last term is constant. Proof: Not so simple.

We have to study sequences u^h such that I^h(u^h) ≤ Ch⁴ and prove some compactness. We first prove that their gradients are locally close to Q(x̄) + x₃B(x̄)

$$\frac{1}{h} \int_{\Omega^h} |\nabla u^h(x) - R^h(\bar{x})(Q(\bar{x}) + x_3 B(\bar{x}))|^2 \, dx \le Ch^4 \tag{2}$$

and the variation of $R^h(\bar{x})$ is controlled

$$\int_{\omega} |\nabla R^h(\bar{x})|^2 \, d\bar{x} \le Ch^2. \tag{3}$$

Then, we prove existence of constants rotations R^h and vectors c^h such that y^h = R^hu^h - c^h enjoys the above properties.

(bis) Limit energy given by

$$\begin{aligned}
 & \frac{\partial \alpha u_{\beta}^{2} + \partial \mu u_{\alpha}^{2} + \partial \alpha u_{3}^{2} \partial \mu u_{\beta}^{2}}{2} \\
 & \mathcal{H}_{VK}(u^{1}, e^{2}) &= 2 \int_{\omega} W_{2} \left(\bar{x}, e^{2} + \frac{1}{2} (\bar{\nabla} u^{1})^{T} \bar{\nabla} u^{1} + \frac{1}{4!} \bar{\nabla} b^{T} \bar{\nabla} b \right) \\
 &+ \frac{1}{3!} \int_{\omega} W_{2} \left(\bar{x}, \bar{\nabla} \varphi^{T} \bar{\nabla} p^{1} + (\bar{\nabla} u^{1})^{T} \bar{\nabla} b \right) \\
 &+ \frac{2}{6!} \int_{\omega} W_{2} \left(\bar{x}, \operatorname{sym}(\bar{\nabla} \varphi^{T} \bar{\nabla} d) + \bar{\nabla} b^{T} \bar{\nabla} b \right)
 \end{aligned}$$

where $p^1 = p^1(u^3)$ (usual case $p^1 = (-\partial_{\alpha} u_3^1, 0)$.

The third term is constant determined by the previous steps and

$$\operatorname{sym}(\bar{\nabla}\phi^{T}\bar{\nabla}d + \bar{\nabla}b^{T}\bar{\nabla}b) = \begin{pmatrix} \mathscr{R}_{1313} & \mathscr{R}_{1323} \\ \mathscr{R}_{1323} & \mathscr{R}_{2323} \end{pmatrix}$$

Therefore, the third term is 0 iff $\Re = 0$, *i.e* the 3d metric is flat. All minima including those of the 3d-problem are 0.

An example: $G(x', x_3) = diag(1, 1, \lambda(x'))$

(i) G is immersible in \mathbb{R}^3 if and only if

$$M_{\lambda} = \nabla^2 \lambda - \frac{1}{2\lambda} \nabla \lambda \otimes \nabla \lambda \equiv 0$$
 in ω ,

(ii) The Γ -limit energy functional $I_{\nu K}$ becomes

$$\begin{aligned} \forall w \in W^{1,2}(\omega, \mathbb{R}^2), \quad \forall v \in W^{2,2}(\omega, \mathbb{R}), \\ I_{vK}v, w) &= 2 \int_{\Omega} W_2(\operatorname{sym} \nabla w + \frac{1}{2} \nabla v \otimes \nabla v + \frac{1}{96\lambda} \nabla \lambda \otimes \nabla \lambda) \, \mathrm{d}x' \\ &+ \frac{1}{3!} \int_{\Omega} W_2(\sqrt{\lambda} \nabla^2 v) + \frac{1}{2 \times 6!} \int_{\Omega} W_2(M_{\lambda}) \, \mathrm{d}x', \end{aligned}$$

where W_2 is independent of x'.